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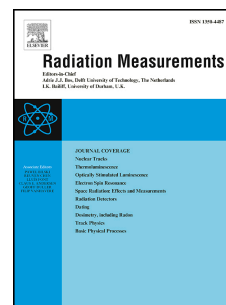
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**Optical determination of the width of the band-tail states, and the excited state and ground state energies of the principal dosimetric trap in feldspar**

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**Abstract**

We constrain parameters that determine thermal stability of the infrared stimulated luminescence (IRSL) signal in a suite of 13 compositionally different feldspar samples by optical probing. We focus specifically on the excited and ground state of the principal trap and the width of the sub-conduction band-tail states. Excitation spectra measured at room temperature result in approximate trap depth of about 2.04 eV and the excited state energy at  $1.44 \pm 0.02$  eV, irrespective of feldspar composition for the sample's measured here. Fitting the non-resonant rising continuum of the excitation spectra suggests that the width of the band-tail states accessible from the ground state of the trap ( $\Delta E$ ) ranges from 0.21 to

0.47 eV at room temperature between the different samples. Photoluminescence measurements are used to constrain the full sub-conduction band-tail width (Urbach width,  $E_u$ ) using the excitation-energy-dependent emission (EDE), resulting in values ranging from 0.26 to 0.81 eV. While the depth of the principal trap and its main excited state seem to be independent of feldspar composition, the difference between  $\Delta E$  and  $E_u$  seems to be related to sample K-content.

## Keywords

Trap depth; IR resonance; band-tail states; excitation spectrum; emission spectrum; excitation-energy-dependent emission

## 1 Introduction

Feldspar is widely used as a chronometer in luminescence dating. The development of post-infrared infrared-stimulated luminescence (post-IR IRSL) protocols (Thomsen et al., 2008; Thiel et al., 2011; Buylaert et al., 2012), which reduce the effects of athermal signal loss (anomalous fading, Wintle, 1973) have increased the use of feldspar in many luminescence studies. Whilst anomalous fading still poses a challenge for some samples, the infrared-stimulated luminescence (IRSL) of feldspar offers a number of advantages over quartz optically stimulated luminescence (OSL), including higher signal sensitivity and saturation dose. Furthermore, the newly developed infrared-photoluminescence (IR-PL) method (Prasad et al., 2017) can be used to date feldspar non-destructively, i.e. without depleting the trapped electron population, opening a range of new research opportunities. The recent methodological developments in feldspar luminescence dating have been paired with an increased understanding of luminescence recombination processes within feldspar

minerals (e.g. Jain and Ankjærgaard, 2011) and several models have been developed to describe the thermal stability of feldspar IRSL (e.g. Jain et al., 2012; 2015; Li and Li, 2013; Guralnik et al., 2015; Lambert et al., In Review).

In luminescence dating applications, provided that the specific luminescence signal under investigation is sufficiently stable relative to the timescale being dated (Aitken, 1985), exact knowledge of the kinetic parameters that govern the signal's stability is not necessary. However, luminescence is increasingly being used to determine rock thermal histories (e.g. Guralnik et al., 2015; King et al., 2016) and such applications require accurate constraint of the parameters that govern thermal decay, in order to extrapolate the cooling history of bedrock over geological timescales. Such extrapolations potentially result in the introduction of large uncertainties on past rock temperatures (Yukihara et al., In Press). Requirement of specialised instrumentation has meant that comparatively few studies have sought to directly characterise the key physical parameters associated with feldspar IRSL thermal decay. The focus of this study is to use optical methods to determine (i) the trap depth and (ii) the excited state energy of the principal trap, and (iii) the width of the sub-conduction band-tail states, in a suite of feldspar samples of different chemical compositions. Analysing a range of different feldspars may provide insights into the cause of the differing thermal stability observed between feldspars that are chemically different (e.g. Tsukamoto et al., 2012).

## **2 Previous investigations of luminescence kinetics of feldspar**

Feldspars are wide band-gap ( $\sim 7.7$  eV) alumino-silicates (Malins et al., 2004), and defects in their crystal lattice allow storage, transport and recombination of charge carriers (Fig. 1). Electrons get trapped in defects (trapping centres) following excitation by ionising

radiation. These trapping centres are able to store electrons over geological time, making trapped-charge dating techniques (luminescence and electron spin resonance dating) possible (Krbetschek et al., 1997; Ikeya, 1993). Electrons can be released from a trap when they receive sufficient energy, either from optical and/or thermal excitation (e.g. Aitken, 1985; Huntley et al., 1985) and when the released electron recombines with a hole (luminescence centre), a luminescence signal is produced (Fig. 1). The thermal stability of a luminescence signal, and thus its validity as a chronometer is at least partly determined by the electron trap depth of the signal under investigation.

## **2.1 The ground and excited state of the IRSL (principal) trap**

Constraining the depth of the principal trap infeldspar has been the subject of fundamental research since the mid-1980s (e.g. Stickertsson, 1985; Hütt et al., 1988) however the exact defect associated with feldspar IRSL remains unclear. Initial research focussed on defining optimal stimulation and emission-detection wavelengths for sediment dating applications (e.g. Hütt et al., 1988; Bailiff and Poolton, 1991; Krbetschek et al., 1996; summary given in Krbetschek et al., 1997) whilst more recent studies have sought to identify the trap depth using the relationship between thermoluminescence (TL) emission peaks and IRSL signals (e.g. Murray et al., 2009; Tsukamoto et al., 2012), or isothermal holding experiments coupled with numerical models (e.g. Guralnik et al., 2015). It has been widely debated whether the luminescence of feldspar is derived from a single trap or multiple traps (e.g. Clark & Sanderson, 1994; McKeever et al., 1997; Baril & Huntley, 2003; Murray et al., 2009; Jain and Ankjærgaard, 2011), although spectroscopic evidence suggests that the same electron trapping centre participates in IRSL and post-IR IRSL signals (Andersen et al., 2012).

The trap depth (with respect to the edge of the conduction band) of a charge trapping centre is determined from the energy required to ionise it. The thermal activation energy required to evict charge from a trap is typically lower (e.g. Stickertsson, 1985; Hütt et al., 1988; Brown and Rhodes, 2017) than the optical energy (e.g. Kars et al., 2013). Hütt et al. (1988) initially constrained the optical activation energy required to release electrons from traps in feldspar, and thus the optical trap depth. They recorded a rising continuum (continuous increase in emission intensity with increasing excitation energy) in the excitation spectra of an irradiated microcline until ~2.25 eV, where their curves peaked (note that the following decrease in signal intensity reported in that study is an instrumental artefact). Poolton et al. (2002a, b) estimated the optical trap depths of NaAlSi<sub>3</sub>O<sub>8</sub> and KAlSi<sub>3</sub>O<sub>8</sub> as 1.97 eV and 1.99 eV, respectively, assuming a hydrogenic model, and validated their results against the experimental data of Bailiff and Poolton (1991) and Spooner (1994). Based on the thermal dependence of time-resolved (TR) green-light stimulated luminescence, Jain and Ankjærgaard (2011) argued that the trap depth of their orthoclase sample was likely to be >2.4 eV (see green arrow in Fig. 1A). Most recently Kars et al. (2013) determined the optical trap depths of NaAlSi<sub>3</sub>O<sub>8</sub> (2.1 eV) and KAlSi<sub>3</sub>O<sub>8</sub> (≥2.5 eV) by fitting excitation spectra measured at 10 K following Bøtter-Jensen et al. (2003).

The excited state energy of the principal trap can also be inferred from excitation spectra. Feldspar show a resonance peak at ~1.45 eV (e.g. Hütt et al., 1988; Baril and Huntley, 2003), with reported full width at half maximum (FWHM) values of 0.12 to 0.13 eV (Godfrey-Smith and Cada, 1996) and 0.11 to 0.21 eV (Baril and Huntley, 2003). Some studies have reported multiple IR excitation peaks (e.g. Hütt et al., 1988; Bailiff and Poolton, 1991; Clark and Sanderson, 1994) and Bøtter-Jensen et al. (1994) observed composition dependent

variations in the shape and position of the resonance peak for plagioclase, albite and potassium-rich feldspar.

## 2.2 Sub-conduction band-tail states

Whilst IR (~1.4 eV) stimulation excites electrons from the ground state to the excited state, it is insufficient for electrons to access the conduction band directly. Poolton et al. (2002a, b) suggested that sub-conduction band-tail states are present within feldspar (Fig. 1A, B), and that they play a key role in electron-hole recombination processes. Band-tail states arise due to imperfections and are distributed in energy levels above the valence band and below the conduction band. Charge in the band-tail states can migrate by hopping (e.g. Morthekai et al. 2012).

Poolton et al. (2009) used photo-transferred OSL (synchrotron photons of energies of 5.5 to 9.0 eV) measured at 10 K to determine the full sub-conduction band-tail width (i.e. Urbach width,  $E_u$ ) for a NaAlSi<sub>3</sub>O<sub>8</sub> (R27) and KAlSi<sub>3</sub>O<sub>8</sub> (R28) sample, obtaining values of 0.28 eV and 0.67 eV respectively. Fitting of energy-resolved stimulation spectra of the same samples (measured at 10 K), but after x-ray irradiation yielded values of 0.32 eV (NaAlSi<sub>3</sub>O<sub>8</sub>, R27) and 0.54 eV (KAlSi<sub>3</sub>O<sub>8</sub>, R28). Poolton et al. (2009) also determined the width of the band-tail states, which are accessible from the ground state of the defect (referred to as  $\Delta E$ ), by fitting the non-resonant rising continuum excitation spectrum of the same samples at 300, 200, 100 and 10 K. They found a decrease in  $\Delta E$  at lower temperatures (from 0.12 to 0.09 eV for NaAlSi<sub>3</sub>O<sub>8</sub> and from 0.30 to 0.16 eV for KAlSi<sub>3</sub>O<sub>8</sub>). Based on this, Poolton et al. (2009) identified two different, temperature dependent, processes governing IRSL in feldspar: (i) a thermally assisted hopping process at temperatures >100 K, and (ii) an athermal process, which disables hopping, although tunnelling remains possible (either via tail-to tail, or tail-to-



recombination centre). Whilst at temperatures  $> 100$  K (i.e. with thermal assistance) a large proportion of band-tail states are accessed resulting in a large  $\Delta E$ , at temperatures  $< 100$  K,  $\Delta E$  is comparably small and invariant with temperature. Kars et al. (2013) used the same approach to determine  $\Delta E$  for the same samples examined by Poolton et al. (2009) and calculated values of 0.18 (R27) and 0.23 eV (R28) for measurements made at 10 K.

More recently Prasad et al. (2016) developed a new method of determining the Urbach width (i.e. the width of the sub-conduction band-tail states, referred to as  $E_u$ ). Using a photoluminescence (PL) emission in the green-orange, they observed an excitation-energy-dependent emission (EDE). The emission peak position shifted linearly, while the peak intensity varied exponentially with a change in the excitation energy; this effect was attributed to the exponentially distributed density of the sub-conduction band-tail states. High excitation energies are able to access band-tail states closer to the conduction band edge, where the density of states is highest and emission intensities are correspondingly larger. Further from the conduction band edge, the density of states is lower and excitation to this part of the band-tail (i.e. using a lower energy excitation) results in a lower emission intensity. Repeating the experiment at 7, 100 and 295 K yielded PL emission spectra that were broadly similar, supporting the hypothesis that the emissions reflect a continuous distribution of the band-tail states. Using this approach, Prasad et al. (2016) obtained  $E_u$  of 0.32 eV for the oligoclase sample investigated. Prasad and Jain (2018) present and discuss a detailed model of the EDE and its application for measuring the band-tail width. Their measurements at 295 K resulted in  $E_u$  values of 0.29 to 0.51 eV. EDE at cryogenic temperatures (7 K) and room temperature (295 K) yielded similar results for their perthitic sample (R58) of 0.33 eV. However, R28 exhibited different band-tail widths at different measurement temperatures: 0.29 eV (295 K) compared to 0.41 eV (7 K). Prasad and Jain

(2018) attribute these differences to an increase in retrapping efficiency, from the lower band-tail states at cryogenic temperatures, related to mobility, which is a function of the net density of states.

As is clear from the discussion above, while there are several articles characterising the excited state energy of the principal trap, only few studies have reported the trap depth and the band tail width. In this contribution, our focus is to characterise the ground and excited state energies of the principal trap and the width of the band tail states using optical methods, in feldspars of different chemical compositions. Specifically, we use excitation spectra to explore (i) the excited and ground state of the IRSL trap and (ii) use the approach of Poolton et al. (2009) and Kars et al. (2013) to fit the rising continuum of the same spectra to indirectly estimate the trap depth and the sub-conduction band-tail states accessible from the ground state of the trap ( $\Delta E$ ). Finally, (iii) following Prasad et al. (2016) and Prasad and Jain (2018), we use EDE to directly measure the width of the sub-conduction band-tail states (i.e. the Urbach width,  $E_u$ ). Characterising the energy levels of the principal trap and the width of the band tail states is necessary for understanding the thermal stability of IRSL in different feldspar samples.

### 3 Material and methods

#### 3.1 Samples

The suite of feldspars selected for this study are summarised in Table 1 and Figure 2. Through characterising samples with different chemical compositions, we can contrast the kinetic parameters of different feldspar with chemical control. We investigated five museum specimens, two sediment samples from Japan (Riedesel et al., 2018, for JSH1-13), and seven bedrock samples from the KTB borehole in Germany, the Japanese Alps, Namche Barwa in

the Himalaya, and the Mont Blanc massif, some of which have been studied previously (Guralnik et al., 2015; King et al., 2016; Lambert et al., In Review). The majority of the samples are part of the alkali feldspar solid solution, although one is an alkali feldspar end-member (Cleavelandite), one is a ternary feldspar (KTB-383-C) and one (DK4) is a plagioclase. Only phase-pure K-feldspar is stable at low (<300 °C) temperature and hence we assume that the alkali feldspars are perthites, comprising intergrowths of K-feldspar (either orthoclase or intermediate microcline) and intermediate albite. One sample, R1-11A, is a microcline cryptoperthite in which the K- and Na-feldspars are coherently intergrown on nm-scales and in which the interfaces between the two phases comprise a significant proportion of the volume of the mineral. F1 is an International Atomic Energy Agency (IAEA AQCS) secondary reference feldspar (IAEA, 1999).

### 3.2 Instrumentation

Measurements were made using the Risø station for Cryogenic Luminescence Research (COLUR; Prasad et al., 2016, 2017), at the Centre for Nuclear Technology, Technical University of Denmark, Risø Campus, Roskilde, Denmark. The measurements were made using a Tungsten halogen lamp, for excitation between 500 and 1000 nm, and a CW Xenon lamp for excitation between 370 and 560 nm. Excitation wavelengths were selected using a double-grating Czerny-Turner monochromator and excitation spectra were recorded through a U340 filter with a UV-sensitive photomultiplier tube (PMT). Emission spectra were detected using the UV-sensitive PMT together with the monochromator. All spectra were corrected for excitation light flux and system response; excitation spectra were converted into photons/s/unit energy interval. Sample material was mounted on the sample holder within COLUR using carbon tape. The sample holder of COLUR comprises the cold finger of a

closed loop He-cryostat that allows temperature dependent measurements to be performed.

A blank aliquot of carbon tape was measured at 300 K to quantify any background signal.

### **3.3 Excitation spectra: Determining trap depth, band-tail width and the excited state of the IRSL (principal) trap**

A large aliquot of sand-sized sample material or a single crystal (~2 mm diameter, < 1 mm thick, sample dependent, see Table 1) were beta irradiated to deposit 500 Gy using a  $\text{Sr}^{90}/\text{Y}^{90}$  beta source attached to a Risø TL/OSL Reader (TL-DA-20). Prior to mounting in the COLUR sample chamber, the samples were preheated for 60 s at 250 °C to remove trapped charge which is thermally unstable over laboratory timescales. Excitation spectra were measured from 500 to 1000 nm, using a step size of 5 nm, a bandpass of 10 nm and an integration time of 1 s. We tested the effect of measurement temperature on selected sample F1, which was measured at 7 K and 300 K; all other measurements were made at 300 K.

### **3.4 Emission spectra: Determining the band-tail width using the excitation-energy-dependent emission (EDE)**

The width of the full sub-conduction band-tail states ( $E_u$ ) of the different samples were determined following Prasad et al. (2016), using fresh, non-irradiated large aliquots of sand-sized samples or single crystals (~2 mm diameter). Samples were mounted on the sample holder in the measurement chamber and excited from 370 to 520 nm using 5, 10 or 20 nm steps, dependent on samples brightness. Emission spectra were recorded with a 1 nm wavelength-scanning interval in the range of 390 to 540 nm, using a 2 nm bandpass. The integration time was either 1 or 10 s, depending on sample brightness. At least two

measurements were repeated for each sample within each measurement cycle to confirm measurement reproducibility, and two samples (NB120 and MBT-I-2430) were measured twice. The results of these repeat measurements were consistent with each other.

## 4 Results

### 4.1 Excitation spectra: Determining trap depth, accessible band-tail width and the excited state of the IRSL trap

The carbon tape substrate (blank measurement) had no detectable emission at any excitation wavelength. The excitation spectrum of each sample showed two distinct features (see examples in Fig. 3B); a resonance peak centred at  $\sim 1.45$  eV in the low energy region (1.24 to 1.70 eV), followed by a monotonic rise to the maximum excitation energy of 2.48 eV measured here. These results are qualitatively consistent with previous studies (e.g., Baril and Huntley, 2003).

The luminescence intensities and the contrast between the two features vary between the different samples (Fig. 3A, B, Figs. S-4-S-7). Cleavelandite (albite) shows the brightest signal intensity whereas plagioclase sample DK4 has very dim luminescence. For all samples except Cleavelandite, F1 and KTB-383-C, the resonant luminescence emission intensity is stronger than that at the highest excitation energy. In addition to the IR resonance peak, MBT-I-2430 shows a broad absorption peak centred at  $\sim 2.1$  eV, which is superimposed on the rising continuum (Fig. 3B).

F1 was measured at 300 and 7 K (Fig. 3A), to investigate the effect of different measurement temperatures. The initial luminescence intensity in response to excitation of 2.25 eV is about an order of magnitude greater at room temperature, relative to

measurement at 7 K, which is significant as some of our samples showed very weak luminescence. The IR resonance peak is absent in the 7 K excitation spectrum (Fig. 3A).

#### 4.1.1 Trap depth

Following Kars et al. (2013) we fitted part of the rising continuum present in the excitation spectra of Cleavelandite, KTB-383-C, NB139, NB120, MBT-F-5704, MBT-I-2430, R1-11A, JSH1-13 and HAM-5 to estimate trap depth using equation [1] (Bøtter-Jensen et al., 2003); the excitation spectra of the other samples were not sufficiently bright.

$$\sigma(E, E_t) \propto \frac{(E - E_t)^{\frac{3}{2}}}{E(E - E_t(1 - \frac{m_0}{m^*}))} \quad [1]$$

where  $\sigma$  is the photo-ionisation cross-section,  $E$  the excitation energy (eV),  $E_t$  the trap depth (eV),  $m^*$  the effective electron mass (fixed at  $m^* = 0.79 m_0$ , Poolton et al., 2001; Kars et al., 2013) and  $m_0$  the electron rest mass. Data were fitted in MATLAB and following Kars et al. (2013), only a part of the rising continuum was fitted (Fig. 4, S-1, -2 and -3), using equation 1 we fitted the luminescence emission intensity for each excitation energy ( $E$ ). Kars et al. (2013) measured excitation spectra up to 2.80 eV, however due to instrumental limitations during the analytical period, the maximum excitation energy explored in this study was 2.48 eV. We tested the influence of fitting different excitation energy ranges on the obtained  $E_t$  to evaluate whether our spectra are sufficiently described (i.e. from 2.10 eV to 2.35 eV in 0.05 eV steps, Table S-1). We recorded an average variance of 7 %, with a trend towards higher  $E_t$  values as the lowest fitted excitation energy increased.

We report values derived from fitting the emission intensity for excitation energies ranging from 2.30 to 2.48 eV and the mean optical trap depth is  $2.04 \pm 0.05$  eV. Sample

MBT-I-2430 exhibit the greatest trap depths of  $2.15 \pm 0.01$  and Cleavelandite has the shallowest trap depth of  $1.99 \pm 0.01$  eV (Table 2). Given that the high-energy plateau predicted by equation 1 is not observed in our experimental data, the model fits are mathematically poorly constrained and these trap depths should be considered as preliminary estimates.

#### 4.1.2 Excited state of the IRSL trap

The IR resonance peak was fitted with a single Gaussian distribution [eq. 2] to characterise the excited state of the IRSL trap.

$$f(E) = k * \exp\left(-\frac{1}{2} * \frac{(E-\mu)^2}{\sigma^2}\right) \quad [2]$$

where  $k$  is a pre-exponential factor,  $E$  the corresponding excitation energy,  $\mu$  the mean energy of the Gaussian distribution and  $\sigma$  the standard deviation. All samples showed similar IR resonance with the Gaussian distribution centred between 1.41 and 1.46 eV (Fig. 3, Figs. S-4, S-5). The FWHM varied from 0.16 to 0.40 eV; results are given in Table 2.

#### 4.1.3 Estimation of band-tail width

To obtain the width of the band-tail states ( $\Delta E$ ) accessible from the ground state of the trap, following Kars et al. (2013) we fitted the rising continuum from 1.80 to 2.48 eV with equation 3 (Poolton et al., 2009).

$$OSL(E) \propto \exp\left(\frac{E-E_t}{\Delta E}\right) \quad [3]$$

where  $E$  is the excitation energy,  $E_t$  the trap depth and  $\Delta E$  the band-tail width accessible from the ground state of the trap. The broad absorption peak at  $\sim 2.1$  eV superimposed on the rising continuum of sample MBT-I-2430 (Fig. 3B, Fig. S-3C) was excluded from the fit of this sample, so that the final excitation energy range considered was limited to 2.22 to 2.48 eV.

The quality of fit varied between samples. It was not possible to calculate  $\Delta E$  for KRG-16-112 or DK4, and some deviation between the fit and the measured data is recorded for Cleavelandite, JSH1-13, HAM-5 and KTB-383-C in the low excitation energy region (Figs S-6 and S-7). Calculated  $\Delta E$  ranges from 0.22 eV (MBT-I-2430) to 0.47 eV (NB120). Compared to the overall range in  $\Delta E$ , measurements of F1 at 300 K (0.29 eV) and 7 K (0.21 eV) yield similar values.

#### 4.2 Estimation of band-tail width using the excitation-energy-dependent emission (EDE)

The PL emission spectra of the different samples were similar and are characterised by a single emission peak. Reducing the excitation energy systematically causes signal intensities to decrease and the emission peak to become narrower and to shift to lower emission energy. At  $\sim 2.2$  eV excitation, no detectable peak is observed. The area under the PL peak was integrated and an exponential relationship between the PL intensity (integrated PL peak area) and the corresponding excitation energy was observed (Fig. 5; Figs. S-8, S-9). Taking the natural logarithm of the integrated PL peak area enabled us to fit a linear regression between  $\ln(\text{PL})$  and the excitation energy (Fig. 5). The inverse of the slope of the fitted line represents the band-tail width. Based on this relationship and following the approach of Prasad et al. (2016) we estimated  $E_u$  for all samples by calculating the inverse of the slope. This resulted in band-tail width estimates ranging from 0.26 to 0.81 eV for



plagioclase sample DK4 and ternary feldspar sample KTB-383-C respectively, full results are listed in Table 2. No shift in peak position was observed for sample R1-11A making it impossible to determine  $E_u$  for this sample.

## 5 Discussion

Previous studies have indicated that the thermal stability of different feldspars may vary (e.g. Tsukamoto et al., 2012). However, no dependence of either the energy of the trap depth or the excited state of the principal trap are apparent for the feldspars investigated here (Fig. 7). The IR resonance of our samples is in agreement with published values (Baril and Huntley, 2003; Poolton et al., 2009) and has an average peak position of  $1.44 \pm 0.02$  eV and FWHM of 0.19 to 0.40 eV. Bøtter-Jensen et al. (1994) proposed that alkali-feldspar chemical composition influences the IR-resonance. However, no clear relationship between mode IR resonance or FWHM, and sample chemical composition, is apparent for our data (Fig. 7).

The similarity between the different feldspar variants indicates that the defect causing IR absorption is likely to be on the feldspar silicate framework, rather than a substitution or defect on the metal site (the compositions and symmetry of which changes substantially across the feldspar mineral group), which is in agreement with previous work (e.g. Short, 2004). However, a specific defect has not yet been associated with IR absorption in feldspar.

The trap depth is a critical parameter for determining the thermal stability. To illustrate, assuming first order kinetic behaviour (which is rarely the case, cf. Guralnik et al., 2015), an  $E_t$  of 2.0 eV yields a lifetime at 50 °C of 13 Ga whereas an  $E_t$  of 1.90 eV yields a lifetime of only 359 Ma assuming a lattice vibration frequency factor ( $s$ ) of  $3.9 \times 10^{13} s^{-1}$ . In

feldspar, the band-tail width also modifies thermal stability (e.g. Li and Li, 2013) and constraining the trap depth of feldspar is challenging because of the presence of the band-tail states (Poolton et al., 2009; Kars et al., 2013). However, despite this it is encouraging that all of our samples, irrespective of chemical composition, yield a nominal average optical trap depth of 2.04 eV with a relatively small standard deviation of 0.05 eV. Furthermore the individual estimates of trap depth are within uncertainties of one another (Figure. 7; Table 2). These preliminary data suggest that the defect or defects interrogated in the optical probing experiments presented here are part of the feldspar aluminosilicate framework, which is broadly consistent between all feldspar variants. In terms of trap depth calculations, a common value of  $E_t$  may be applicable to feldspar of different compositions, and further investigations using low temperature OSL excitation spectra, and/or newly developed optical probing methods to directly characterise the IRSL trap depth without the effects of competing recombination processes (Prasad et al., 2017), will be useful to investigate this potential in the future.

Regarding the width of the sub-conduction band-tail states, both methods, viz. the excitation spectra measured at 300 K to determine  $\Delta E$  (equation 3) and EDE (Prasad et al., 2016; Prasad and Jain, 2018) used to determine  $E_u$  yield broadly similar results for some samples (Fig. 6).  $\Delta E$  ranged from 0.21 to 0.46 eV and  $E_u$  from 0.26 to 0.81 eV (Figs. 6, 7). These results are within the range of previously published values (Poolton et al., 2009; Prasad and Jain, 2018). The reference feldspars that we investigated (F1, Cleavelandite and DK4; listed from medium to low potassium content), exhibit a decrease in  $E_u$  with reducing K-content from  $0.48 \pm 0.03$  eV (orthoclase content: 61.3%) to  $0.26 \pm 0.02$  eV (orthoclase component: 0.5%). However, this trend is not apparent in the non-reference feldspar samples measured as large aliquots of sand-sized material, which may be a consequence of

signal averaging across grains of different chemical composition. Single-phase feldspars are unlikely in this sand-sized material, and the feldspars under investigation are likely to comprise intergrowths of K-feldspar and Na-feldspar. Consequently the apparent consistency might be related to one phase dominating the response in these experiments.

The relationship between  $E_u$  and  $\Delta E$  is also interesting (Figure 6). Our results suggest that  $\Delta E$  is either consistent with  $E_u$  (five samples) or smaller than  $E_u$  (five samples). This relationship confirms expectation because whereas  $E_u$  (the Urbach band-tail width) mainly maps the excitation efficiency,  $\Delta E$  (the width of the band-tail accessible from the ground state of the trap) depends on both the excitation efficiency and the recombination efficiency, which in turn is a function of the energy of the band tail state (Jain and Ankjær, 2011) and temperature. Interestingly four out of five samples with  $\Delta E < E_u$  have relatively low K contents (F1, JSH1-13, HAM-5, KTB-383-C), potentially indicating reduced recombination efficiency with band-tail energy in K-poor feldspars. Poolton et al. (2009) and Kars et al. (2013) also found that  $\Delta E$  is narrower in  $\text{NaAlSi}_3\text{O}_8$  than in  $\text{KAlSi}_3\text{O}_8$ , providing support for this observation.

Whilst it is important to consider the absolute width of the band-tail states, the absolute density of band-tail states will also impact the thermal stability of luminescence signals in feldspars. This aspect has not been investigated in the present study but should be considered for assessing thermal stability of IRSL.

## 6 Conclusion

In this study we investigated the kinetic parameters of feldspar that are thought to determine thermal decay using optical probing, and attempted to relate these parameters to feldspar chemical composition. We obtained a nominal average estimate for the optical trap

depth of  $2.04 \pm 0.05$  eV and the energy of the excited state of the IRSL trap of  $1.44 \pm 0.02$  eV, with FWHM ranging from 0.16 to 0.40 eV. Detailed analyses of both the full sub-conduction band-tail width ( $E_u$ ) and the sub-conduction band-tail states accessible from the ground state of the trap ( $\Delta E$ ), revealed values ranging from 0.2 to 0.8 eV. All values are consistent with previous studies. Interestingly the difference between  $\Delta E$  and  $E_u$  is greater for samples with low K-content, potentially indicating that recombination efficiency must vary more with band tail energy in K poor feldspars. Based on these findings we tentatively conclude that the defect or the defects leading to IRSL in feldspar are within the alumino-silicate framework of feldspar, which is almost the same for all alkali feldspars and would explain the similarity in our results. In contrast to the trap energy levels (ground and excited state),  $\Delta E$ ,  $E_u$  and the FWHM of the excited state exhibited larger variability between samples; combining these data with numerical models of feldspar thermal decay, and contrasting them with isothermal decay data in future work will provide further insights into which parameters are the key drivers of differences in the thermal stability of feldspars.

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**Figure captions**

**Fig. 1:** Energy band model with density of states, recombination processes and energy transitions of the principal trap in feldspar. A) Distribution of the density of states, Gaussian distributions correspond to different energy states (e.g. the ground state and excited state of the IRSL (principal) trap in feldspar). While the density of sub-conduction band-tail states are thought to be exponentially distributed with energy below the conduction band edge, energy states in the conduction band follow a parabolic shape (modified from Poolton et al., 2002b; Jain & Ankjærgaard, 2011). The full sub-conduction band-tail width (i.e. the Urbach width) is referred to as  $E_u$  (as they are also called Urbach tail), and may extend below the ground state of the principal trap, whereas  $\Delta E$  defines the band-tail states accessible from the ground state of the defect. B) Possible transitions from the ground state of the IRSL trap (modified from Jain & Ankjærgaard, 2011). a) Thermal excitation. b) Optical excitation by different wavelengths; whereas blue light stimulation allows immediate transition into the conduction band, green light results in a sub-conduction band transition and IR stimulation excites the electron to the excited state. c) Recombination through the conduction band. d) Recombination through the band-tail states. When the electron has been excited to the excited state of the trap by IR stimulation, phonon-assisted diffusion (PAD) allows transition through the band-tail states. e) Athermal signal loss due to quantum mechanical tunnelling from the ground state of the trap. f) Retrapping, which may result in a photon with the same energy as the absorbed photon.

**Fig. 2:** Feldspar ternary of chemical composition and origin map for samples in this study.

**Fig. 3:** Example excitation spectra: A) Rising continuum of F1, measured at 7 and 300 K; B) Excitation spectra of Cleavelandite, KRG-16-112 and MBT-I-2430. These spectra are representative for the range of sample brightness and the shape of the IR resonance peak and non-resonant rising continuum. In contrast to the other samples investigated, MBT-I-2430 also shows a peak centred at  $\sim 2.10$  eV superimposed on the rising continuum.

**Fig. 4:** Excitation spectrum of KTB-383-C showing model fits. The IR resonance peak was fitted with a single Gaussian distribution [eq. 2] and the rising continuum was fitted with equation 1 (Bøtter-Jensen et al., 2003) to estimate trap depth and equation 3 (Poolton et al., 2009) to obtain an estimate of the band-tail width accessible from the ground state of the trap ( $\Delta E$ ). Each part of the excitation spectrum was fitted separately.

**Fig. 5:** Photoluminescence emission spectra of the excitation-energy-dependent emission (EDE) and the full band-tail width ( $E_u$ ) calculation for Cleavelandite. A) EDE spectra. B) Normalised EDE spectra. C) Exponential relationship between the integrated area under the emission spectrum and the corresponding excitation energy. The inverse of the slope ( $1/e$ ) gives  $E_u$ .

**Fig. 6:** Comparison of band-tail states accessible from the ground state of the trap ( $\Delta E$ ) and the full band-tail width ( $E_u$ ) calculated from fitting the rising continuum with eq. 3 and by using EDE respectively. The dashed line indicates the 1:1 relationship.

**Fig. 7:** Summary of trap depth, band-tail width and position of the IR resonance peak, obtained from optical measurements. Results of this study are compared to published

research. Baril & Huntley (2003) investigated 19 samples, since all their samples show a main IR-resonance at 1.44–1.45 eV, just a single point represents their data.

**Table 1:** List of samples with chemical composition (determined using the Risø XRF OSL/TL reader attachment, Kook et al., 2012) and geological origin.

ID	Type <sup>1</sup>	Origin	Chemical composition <sup>2</sup> (FS % <sup>3</sup> )			Quartz amount (%)	Geological age
			Or	Ab	An		
NB139	Bedrock	Migmatitic gneiss, Namche Barwa, Nepal	89.3	9.4	1.4	4.2	Proterozoic (Burg et al., 1997)
NB120	Bedrock	Migmatitic gneiss, Namche Barwa, Nepal	86.2	11.0	2.9	5.0	Proterozoic (Burg et al., 1997)
MBT-I-2430	Bedrock	Calc-alkaline granite, Mont Blanc Tunnel, Italy	86.7	12.0	1.3	4.2	~300 Ma (Bussy et al., 1989; Bussy and Von Raumer, 1993)
MBT-F-5704	Bedrock	Calc-alkaline granite, Mont Blanc Tunnel, France	76.3	22.5	1.2	4.6	~300 Ma (Bussy et al., 1989; Bussy and Von Raumer, 1993)
KRG-16-06	Bedrock	Kurobegawa granite, Japanese Alps, Japan	79.4	17.3	3.2	3.1	<10 Ma (Ito et al., 2013)
KGR-16-112	Bedrock	Kurobegawa granite, Japanese Alps, Japan	75.2	21.3	3.5	3.4	<10 Ma (Ito et al., 2013)
R1-11A	Museum	Rapakivi Granite, South Greenland	68.0	22.0	10.0		c. 1750 Ma (Finch and Klein, 1999)
F1	IAEA AQCS Reference Feldspar		61.3	33.4	5.4	0.3	NA
JSH1-13	Sediment	Shirasuka Lowlands, Japan	50.4	37.6	12.1	38.1	NA
HAM-5	Sediment	Lake Hamana, Japan	64.9	31.1	4.0	19.3	NA
KTB-383-C	Bedrock	KTB borehole, Germany	18.2	67.6	14.4	52.8	
Cleavelandite	Museum	Pegmatite, Golonca District, Minas Gerais, Brazil	0.5	99.3	0.2	3.9	650-450 Ma (Cassedanne and Roditi, 1996)
DK4	Museum	Klokken intrusion, Sout Greenland	5.0	36.0	59.0		1.18 Ma

Notes:

<sup>1</sup> Museum specimen are single crystals, the IAEA reference feldspar, feldspars extracted from bedrock and sediment samples were mounted as coarse multiple-grain aliquots

<sup>2</sup> Determined by XRF measurements

<sup>3</sup> Assuming 100 % feldspar

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**Table 2:** Optical kinetics of feldspar samples investigated in this study. The reported uncertainties account only for the fitting, no uncertainties resulting from the instrumentation are added.

Sample <sup>1</sup>	Optical trap depth (eV) <sup>2</sup>	Excited state (IR resonance peak) <sup>3</sup>			Band-tail width ( $\Delta E$ , fitting using eq. 3)		Band-tail width ( $\Delta E$ )
		Mode (eV)	Standard deviation (sigma, eV)	FWHM	$\Delta E$ (eV)	Fitted excitation energy range (eV)	
NB139	2.03 $\pm$ 0.02	1.46	0.09	0.21	0.39 $\pm$ 0.01	1.80 – 2.48	0.37 $\pm$ 0.01
NB120	2.03 $\pm$ 0.04	1.45	0.11	0.26	0.47 $\pm$ 0.02	1.80 – 2.48	0.42 $\pm$ 0.01
MBT-I-2430	2.15 $\pm$ 0.01	1.45	0.07	0.16	0.22 $\pm$ 0.01	2.22 – 2.48	0.32 $\pm$ 0.01
MBT-F-5704	2.05 $\pm$ 0.01	1.44	0.08	0.19	0.31 $\pm$ 0.01	1.80 – 2.48	0.30 $\pm$ 0.03
KRG-16-06	NA	1.44	0.14	0.33	0.46 $\pm$ 0.03	1.80 – 2.48	0.42 $\pm$ 0.01
KRG-16-112	NA	1.41	0.16	0.38	NA	NA	0.45 $\pm$ 0.01
R1-11A	2.02 $\pm$ 0.03	1.43	0.12	0.28	0.38 $\pm$ 0.01	1.80 – 2.48	No shift
F1	NA	1.45	0.10	0.24	0.29 $\pm$ 0.01	1.80 – 2.48	0.48 $\pm$ 0.03
F1 (7 K)	2.10 $\pm$ 0.02	no peak	no peak	no peak	0.21 $\pm$ 0.00	1.80 – 2.48	no peak
JSH1-13	2.01 $\pm$ 0.01	1.43	0.10	0.24	0.36 $\pm$ 0.01	1.80 – 2.48	0.50 $\pm$ 0.01
HAM-5	2.03 $\pm$ 0.01	1.43	0.10	0.24	0.31 $\pm$ 0.00	1.80 – 2.48	0.42 $\pm$ 0.01
KTB-383-C	2.04 $\pm$ 0.01	1.43	0.09	0.21	0.30 $\pm$ 0.00	1.80 – 2.48	0.81 $\pm$ 0.03
Cleavelandite	1.99 $\pm$ 0.01	1.45	0.10	0.24	0.34 $\pm$ 0.01	1.80 – 2.48	0.33 $\pm$ 0.00
DK4	NA	1.41	0.17	0.40	NA	NA	0.26 $\pm$ 0.02

Notes:

<sup>1</sup> Samples are sorted from high to low potassium content, DK4 is a plagioclase.

<sup>2</sup> Excitation energy range used for fitting with equation 1: 2.30 – 2.48 eV

<sup>3</sup> The parameters refer to the fitted Gaussian distribution (eq. 2), not to the peak itself.

## Supplementary Material

**Fig. S-1:** Excitation spectra fitted using eq. 1 (Bøtter-Jensen et al., 2003) to determine trap-depth ( $E_t$ ) for F1, which was measured at 7 K. The solid line shows the fit and the dashed line the extrapolated curve. The intersect of the dashed curve with the x-axis gives the trap depth.

**Fig. S-2:** Excitation spectra fitted using eq. 1 (Bøtter-Jensen et al., 2003) to determine trap-depth ( $E_t$ ) for samples NB139, NB120, MBT-I-2430, MBT-F-5704, R1-11A and JSH1-13, all measured at 300 K. The solid line shows the fit and the dashed line the extrapolated curve. The intersect of the dashed curve with the x-axis gives the trap depth.

**Fig. S-3:** Excitation spectra fitted using eq. 1 (Bøtter-Jensen et al., 2003) to determine trap-depth ( $E_t$ ) for samples HAM-5, KTB-383-C and Cleavelandite, all measured at 300 K. The solid line shows the fit and the dashed line the extrapolated curve. The intersect of the dashed curve with the x-axis gives the trap depth.

**Fig. S-4:** Excitation spectra showing the fit of the IR resonance peak of samples NB139, NB120, MBT-I-2430, MBT-F-5704, KGR-16-06, KRG-16-112, R1-11A and F1 with a single Gaussian distribution [eq. 2].

**Fig. S-5:** Excitation spectra showing the fit of the IR resonance peak of samples JSH1-13, HAM-5, KTB-383-C, Cleavelandite and DK4 with a single Gaussian distribution [eq. 2].

**Fig. S-6:** Excitation spectra of samples NB139, NB120, MBT-I-2430, MBT-F-5704, KRG-16-06, KRG-16-112 and R1-11A fitted with equation 3 (Poolton et al., 2009) to obtain the width of the band-tail accessible from the ground state of the trap ( $\Delta E$ ).

**Fig. S-7:** Excitation spectra of samples F1 (300 K and 7 K), JHS1-13, HAM-5, KTB-383-C, Cleavelandite and DK4 fitted with equation 3 (Poolton et al., 2009) to obtain the width of the band-tail accessible from the ground state of the trap ( $\Delta E$ ).

**Fig. S-8:** Exponential relationship of the integrated area under the PL peak and the corresponding excitation energy of samples NB139, NB120, MBT-I-2430, MBT-F-5704, KRG-16-06 and KRG-16-112. The inverse of the slope gives the band-tail width ( $E_u$ ).

**Fig. S-9:** Exponential relationship of the integrated area under the PL peak and the corresponding excitation energy of samples F1, JHS1-13, HAM-5, KTB-383-C, Cleavelandite and DK4. The inverse of the slope gives the band-tail width ( $E_u$ ).

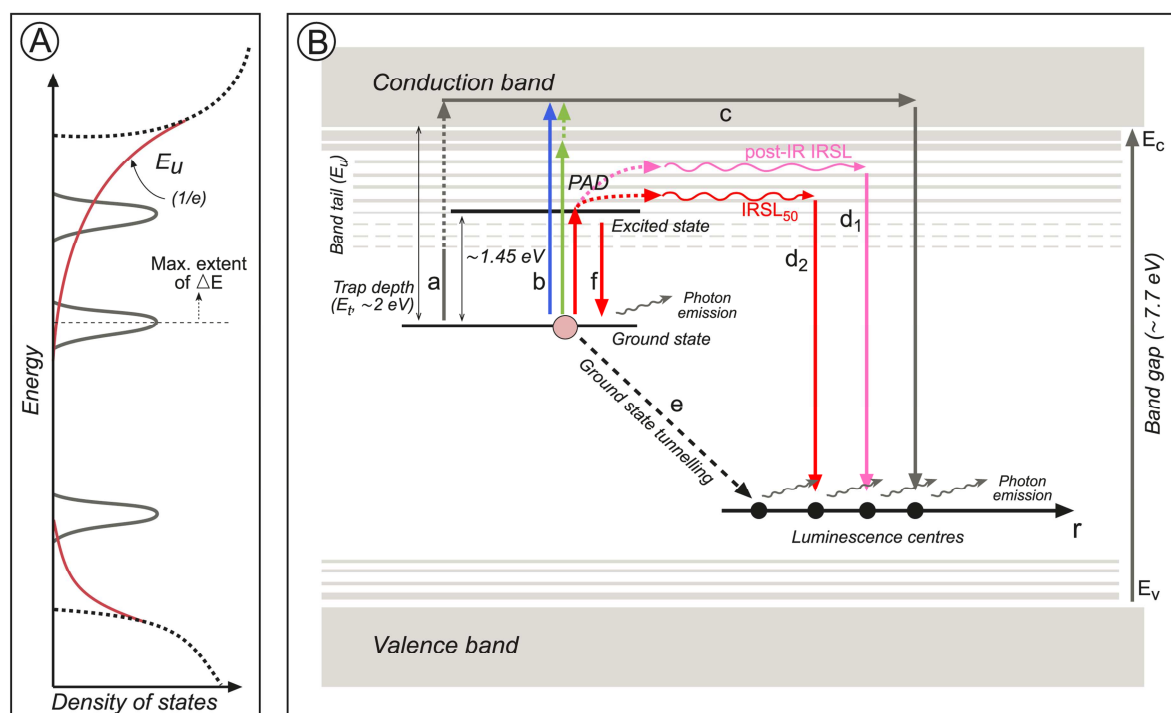
**Table S-1:** Trap depth estimates using equation 1 (Bøtter-Jensen et al., 2003) for different integration ranges.

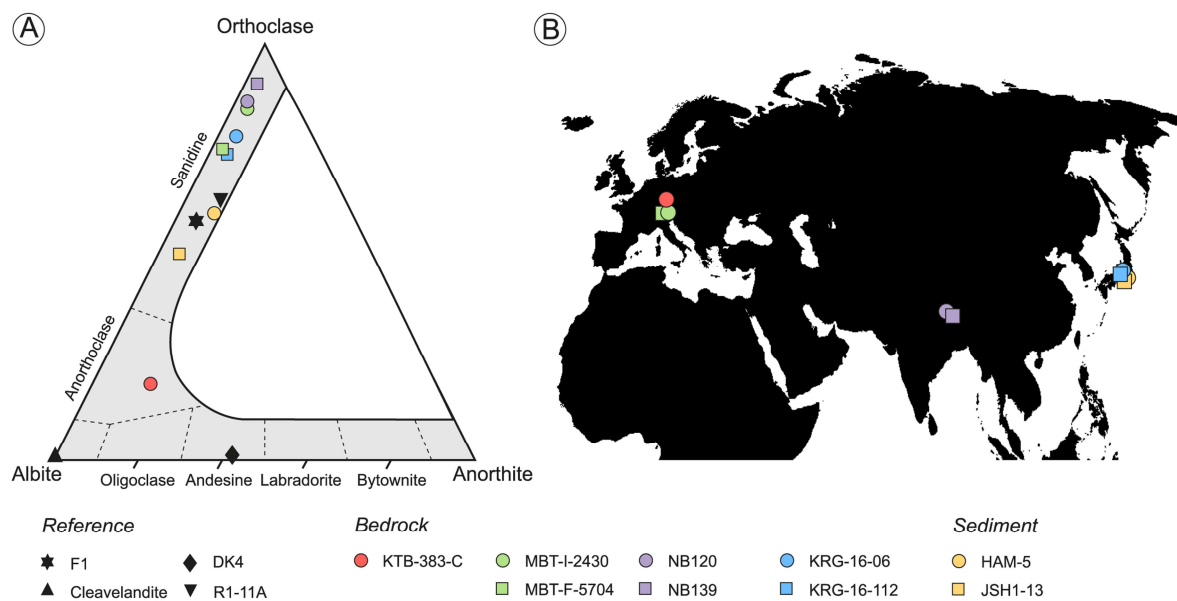
Sample <sup>1</sup>	Optical trap depth (eV)					
	2.10–2.48 eV	2.15–2.48 eV	2.20–2.48 eV	2.25–2.48 eV	2.30–2.48 eV	2.35–2.48 eV
NB139	1.86 ± 0.02	1.90 ± 0.02	1.94 ± 0.02	1.97 ± 0.02	2.03 ± 0.2	2.06 ± 0.03
NB120	1.84 ± 0.03	1.89 ± 0.03	1.95 ± 0.03	2.01 ± 0.03	2.03 ± 0.04	2.09 ± 0.05
MBT-I-2430	1.95 ± 0.02	2.01 ± 0.02	2.07 ± 0.01	2.11 ± 0.03	2.15 ± 0.01	2.19 ± 0.01
MBT-F-5704	1.93 ± 0.01	1.97 ± 0.01	2.00 ± 0.01	2.03 ± 0.01	2.05 ± 0.01	2.07 ± 0.02
R1-11A	1.88 ± 0.02	1.91 ± 0.02	1.96 ± 0.02	1.98 ± 0.03	2.02 ± 0.03	2.06 ± 0.05
F1 <sup>2</sup>	2.02 ± 0.01	2.05 ± 0.01	2.06 ± 0.01	2.08 ± 0.01	2.10 ± 0.02	2.11 ± 0.03
JSH1-13	1.88 ± 0.01	1.90 ± 0.01	1.94 ± 0.01	1.97 ± 0.01	2.01 ± 0.01	2.02 ± 0.01
HAM-5	1.92 ± 0.01	1.95 ± 0.01	1.98 ± 0.01	2.01 ± 0.01	2.03 ± 0.01	2.04 ± 0.01
KTB-383-C	1.93 ± 0.01	1.96 ± 0.01	1.99 ± 0.01	2.02 ± 0.01	2.04 ± 0.01	2.05 ± 0.01
Cleavelandite	1.88 ± 0.01	1.91 ± 0.01	1.94 ± 0.01	1.97 ± 0.01	1.99 ± 0.01	2.01 ± 0.01

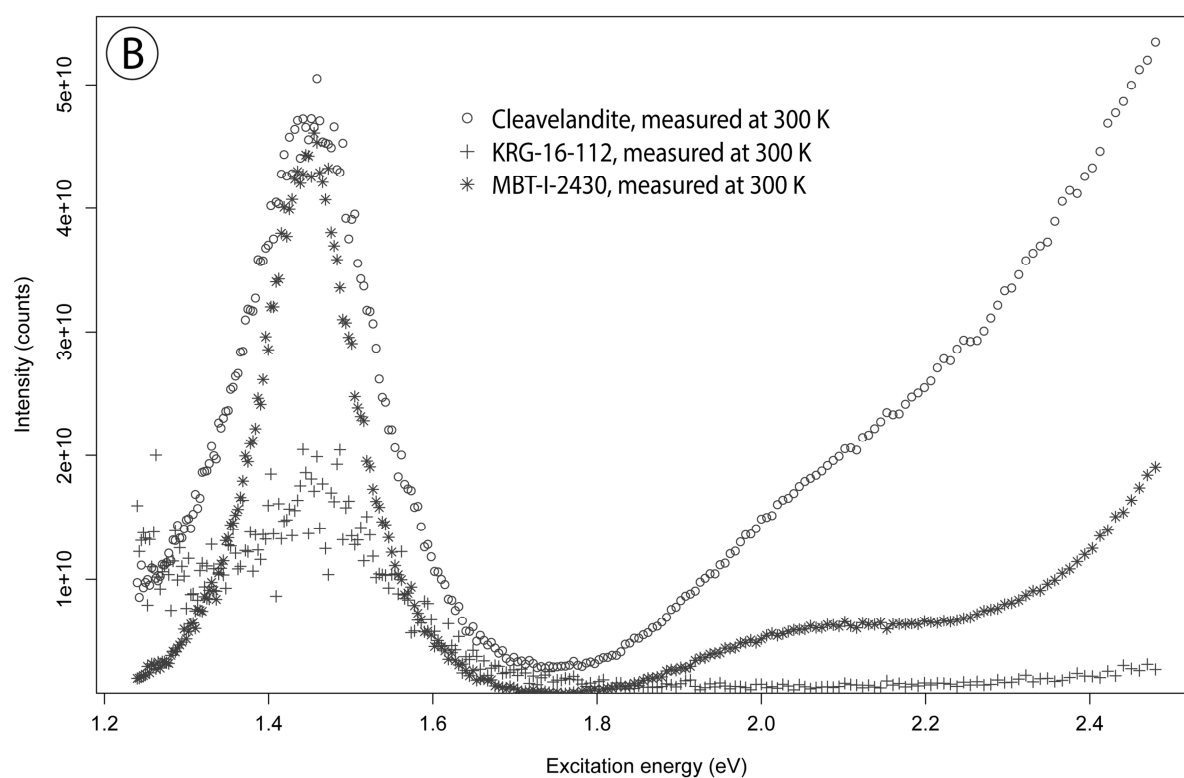
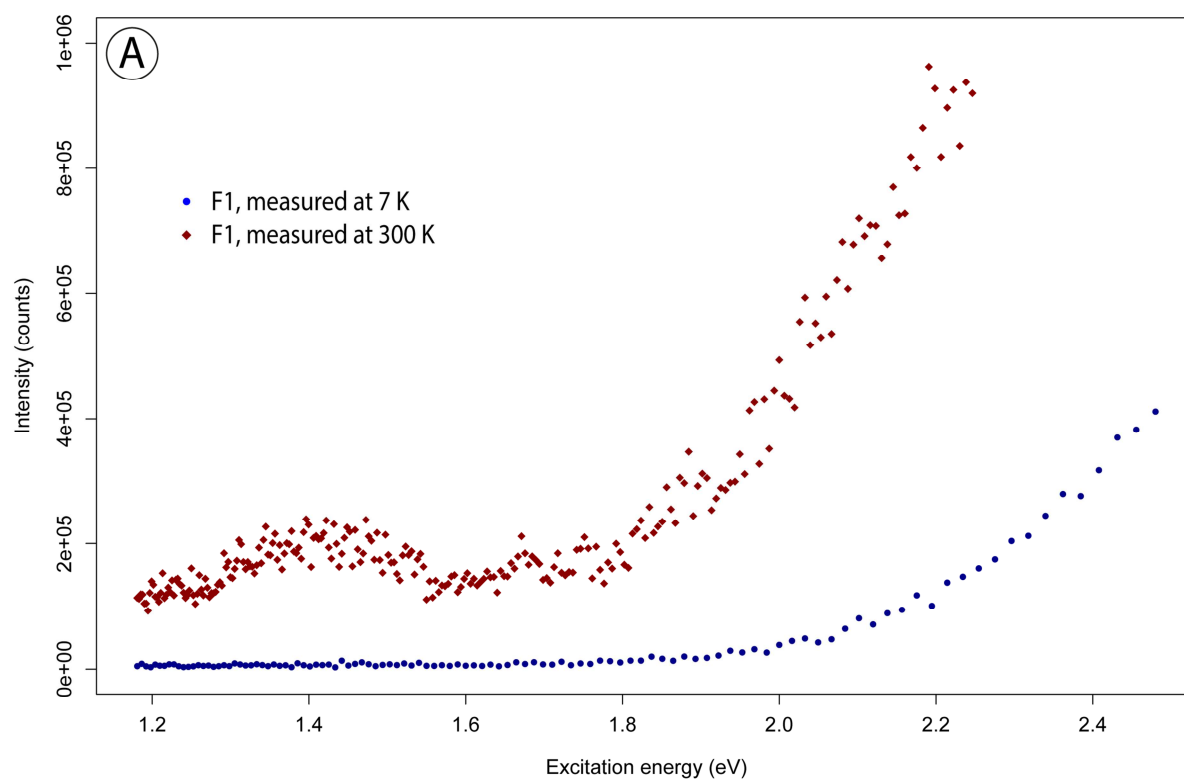
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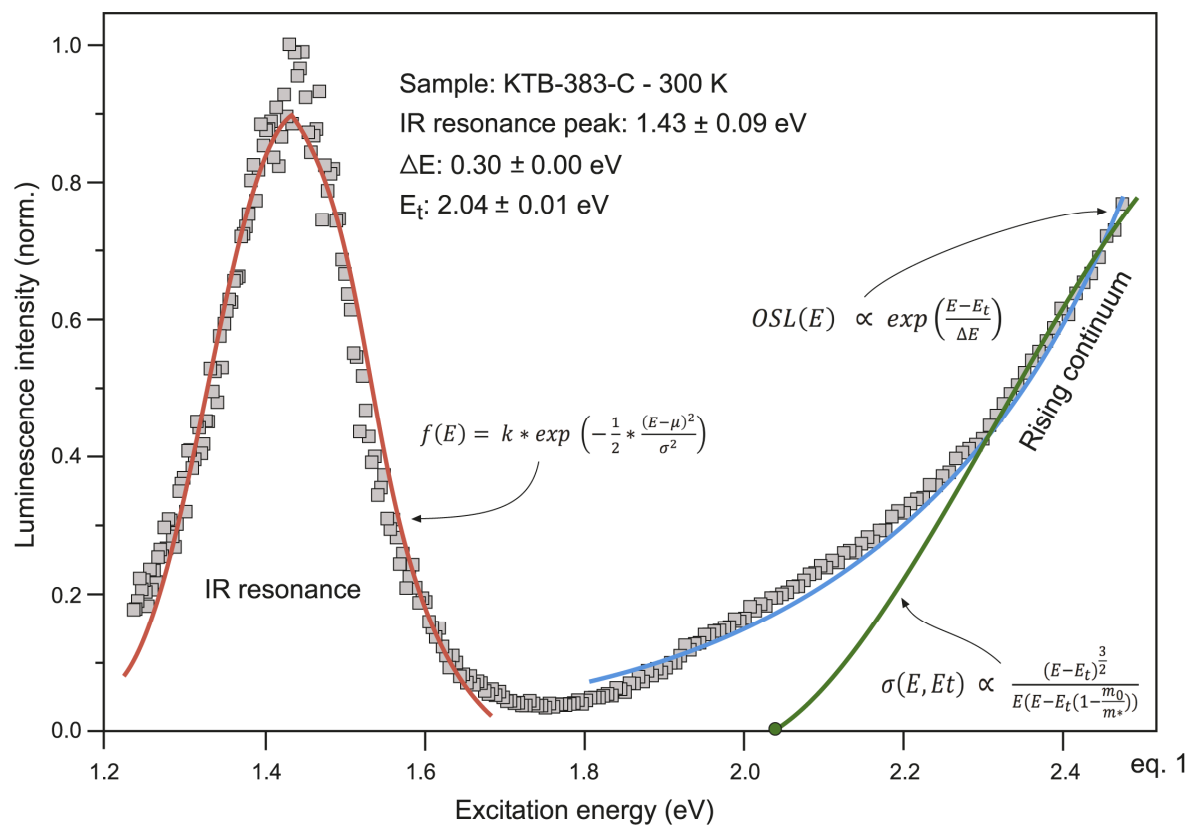
<sup>1</sup> Samples are sorted from high to low K-content.

<sup>2</sup> F1 was measured at 7 K.

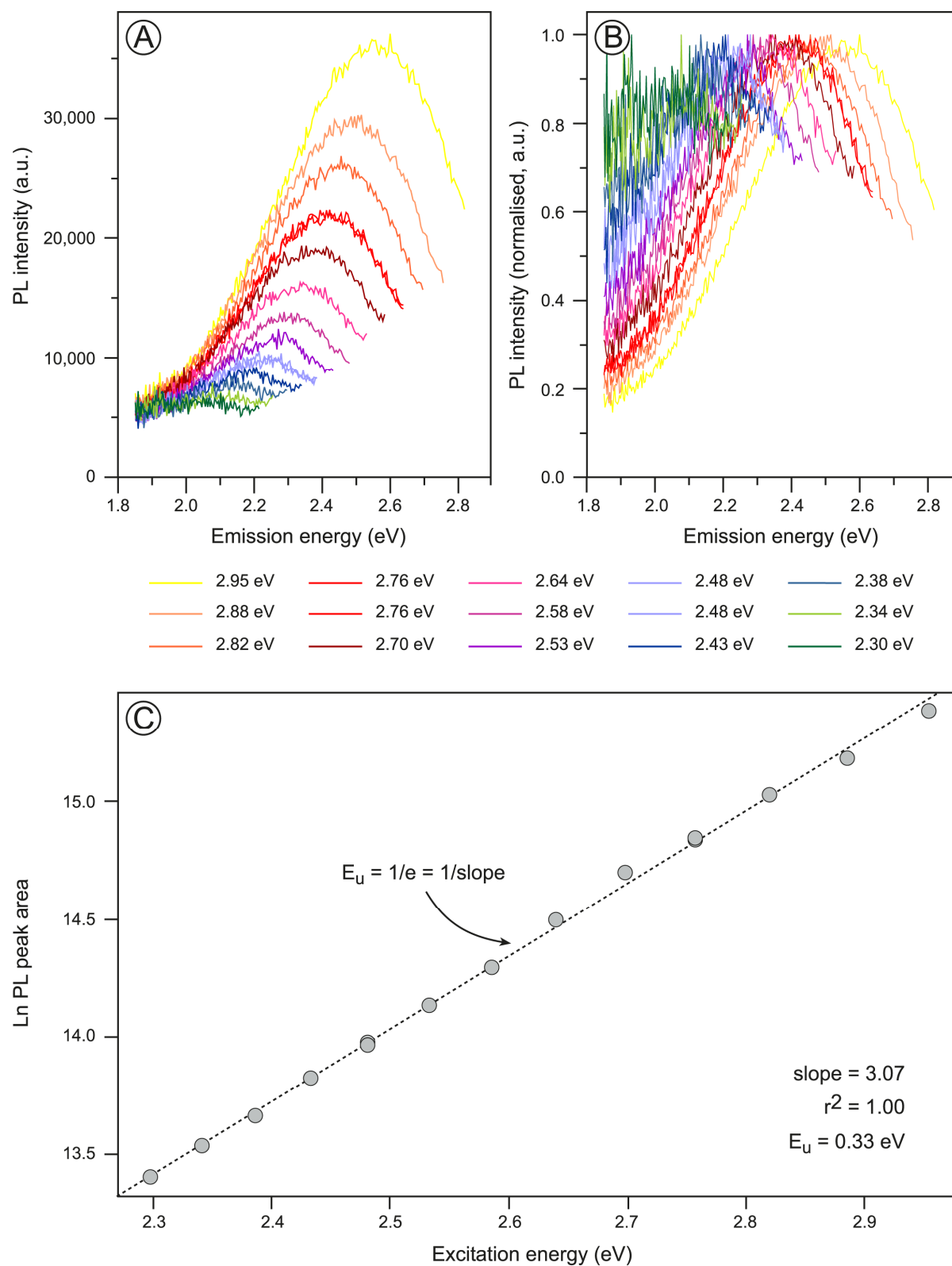


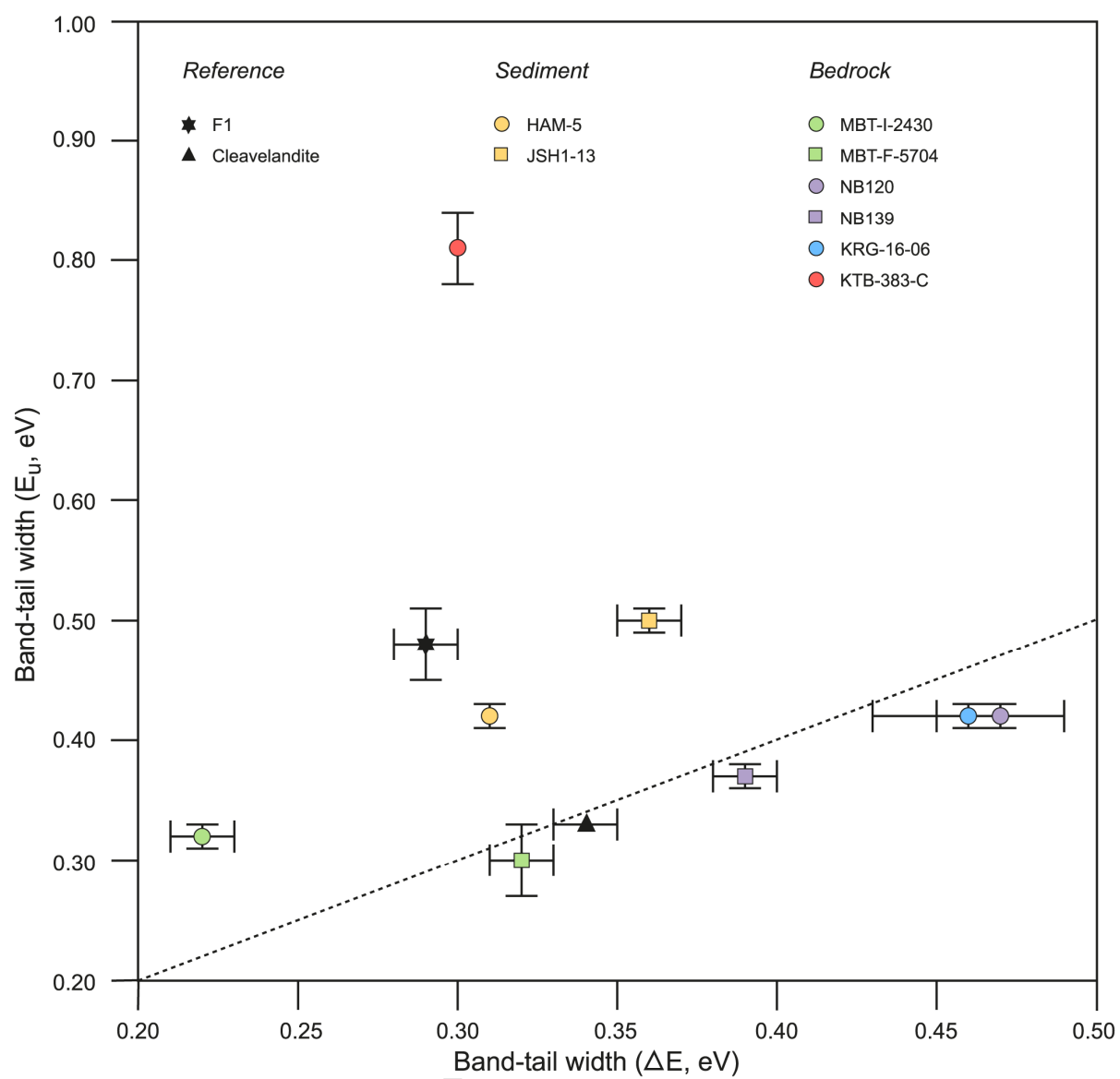


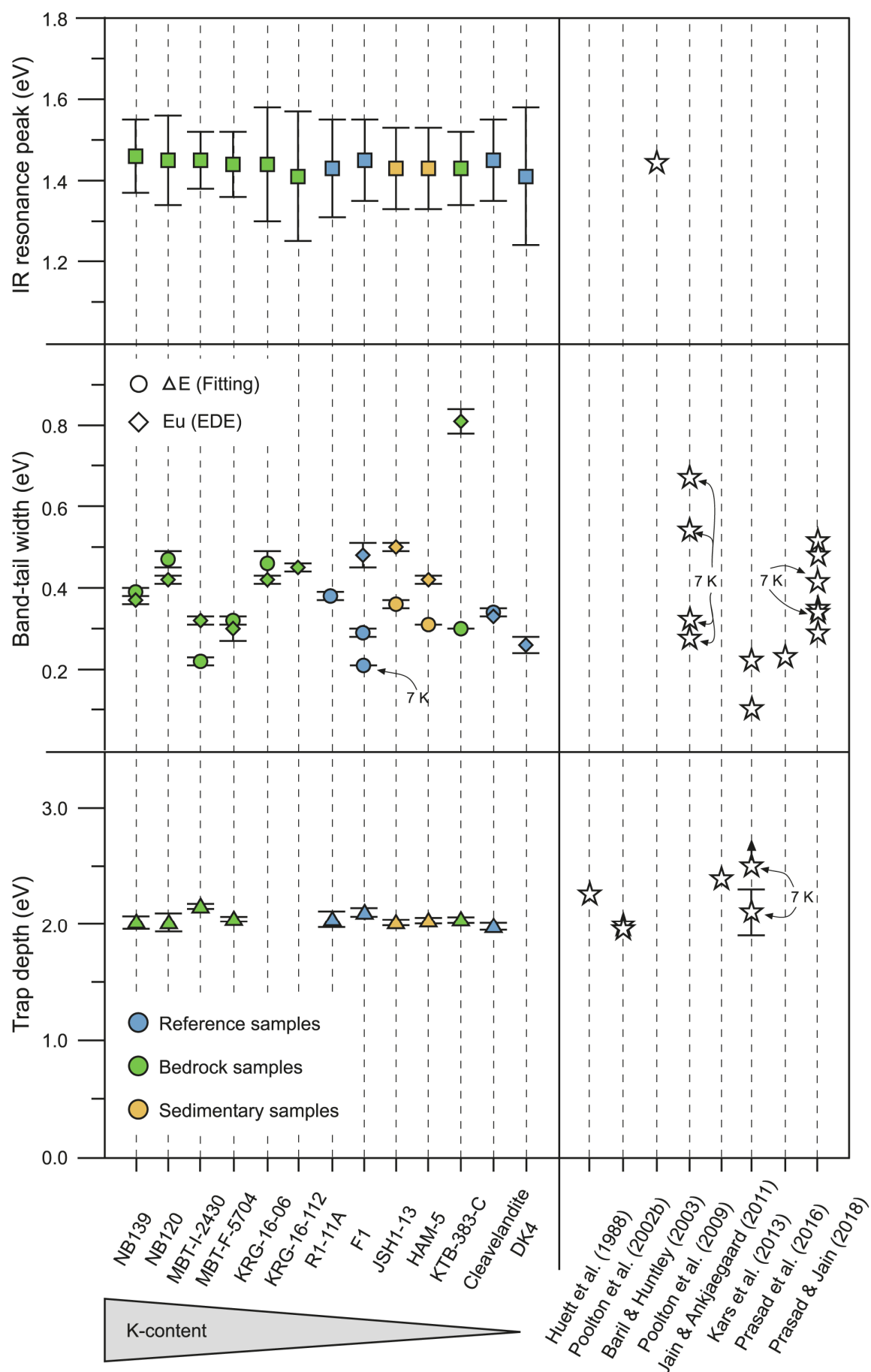












**Highlights**

- Thirteen chemically different feldspars characterised
- Ground and excited state energies of the trap invariant with chemical composition
- Sub-conduction band-tail states width more variable between feldspars